

S to Co coordination in  $\text{Co}_9\text{S}_8$  is that of the  $\text{Rh}(m)$  with the  $\text{S}(f)$ ,  $\text{S}(i)$  and  $\text{S}(j)$  atoms. The average distance for these is 2.33 Å (Geller, 1962*b*) compared with the average tetrahedral Co–S distance, 2.19 Å. Moreover, the relatively longer  $\text{Rh}(m)$ – $\text{Rh}(m)$  than  $\text{Co}(f)$ – $\text{Co}(f)$  distances (Table 4) would imply a relative shortening of  $\text{Rh}(m)$ –S distances. Thus we are able to conclude that the 4-coordination distances are relatively as one would have predicted.

It is possible that the more complex bonding and packing in the  $\text{Rh}_{17}\text{S}_{15}$  results in the shortening of the bonds of the single Rh atom in the unit cell having octahedral coordination. By studying the systems in which Co replaces Rh in  $\text{Rh}_{17}\text{S}_{15}$  and that in which Rh replaces Co in  $\text{Co}_9\text{S}_8$ , we hope to gain some further insight into the bonding and atom sizes as well as effects on the superconductive properties of  $\text{Rh}_{17}\text{S}_{15}$ .

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## The Crystal Structure of the Superconductor $\text{Rh}_{17}\text{S}_{15}$

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Crystals of  $\text{Rh}_{17}\text{S}_{15}$ , isostructural with those of  $\text{Pd}_{17}\text{Se}_{15}$ , belong to one of three space groups  $O_h^1$ – $Pm\bar{3}m$ ,  $T_d^2$ – $P\bar{4}3m$  or  $O^2$ – $P432$ , have a lattice constant of  $9.911 \pm 0.001$  Å, and two formula units per unit cell. The structure of  $\text{Rh}_{17}\text{S}_{15}$  has been refined in  $Pm\bar{3}m$  from selected single-crystal X-ray data. It is shown that there are relatively short Rh–Rh contacts in  $\text{Rh}_{17}\text{S}_{15}$ .

### Introduction

To date, only  $\text{Rh}_{17}\text{S}_{15}$  and  $\text{Pd}_{17}\text{Se}_{15}$  are known to crystallize with the structure to be described in this paper. A detailed account of the structural study of  $\text{Pd}_{17}\text{Se}_{15}$  has already been given (Geller, 1962); it was mentioned that X-ray data from both  $\text{Rh}_{17}\text{S}_{15}$  and  $\text{Pd}_{17}\text{Se}_{15}$  were useful in the solution of the structural problem. The main purpose of the present paper is to report on the structural details of  $\text{Rh}_{17}\text{S}_{15}$  and to point out how they differ from those of  $\text{Pd}_{17}\text{Se}_{15}$ .

The rhodium–sulfur system was first investigated by Juza, Hulsmann, Meisel & Biltz (1935). They found a compound that they designated  $\text{Rh}_9\text{S}_8$ , which appeared from X-ray powder diffraction data to have primitive cubic symmetry. The compound was found to be superconducting by Matthias, Corenzwit & Miller (1954). Recently, Kjekshus (1960) carried out X-ray studies over a temperature range on compounds alleged to be  $\text{Rh}_9\text{S}_8$  and  $\text{Pd}_9\text{Se}_8$ . The structural study has conclusively shown, however, that the correct formulae of the compounds are  $\text{Rh}_{17}\text{S}_{15}$  and  $\text{Pd}_{17}\text{Se}_{15}$ , as shown in the paper on the latter compound.

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### Experimental

The lattice constant obtained from a powder photograph (taken with a 114.6 mm. diameter Norelco Straumanis camera and using Cu *K* radiation) of a specimen with an originally intended composition  $\text{Rh}_9\text{S}_8$  is  $9.911 \pm 0.001$  Å. This value probably represents the lattice constant of  $\text{Rh}_{17}\text{S}_{15}$  well within experimental error. A compound  $\text{Rh}_9\text{S}_8$  would contain 78.31% Rh by weight while  $\text{Rh}_{17}\text{S}_{15}$  contains 78.44% Rh by weight; the difference is small. Furthermore a specimen made up with intended composition  $\text{Rh}_8\text{S}_7$  (78.58% Rh) contained elementary rhodium as shown by an X-ray photograph. Excess sulfur would be difficult to see in the ' $\text{Rh}_9\text{S}_8$ .' The phase appears to be fairly sharply defined in the Rh–S system, but even if there were some solution of sulfur in the  $\text{Rh}_{17}\text{S}_{15}$ , there would probably not be a significant difference in lattice constant between ' $\text{Rh}_9\text{S}_8$ ' and  $\text{Rh}_{17}\text{S}_{15}$ .

With two  $\text{Rh}_{17}\text{S}_{15}$  in the unit cell, the X-ray density 7.60 g.cm.<sup>-3</sup> is in good agreement with the value 7.68 g.cm.<sup>-3</sup> obtained pycnometrically on an ' $\text{Rh}_9\text{S}_8$ ' specimen.

An essentially\* single crystal of irregular shape and less than 0.2 mm. each dimension, was photographed with a Buerger precession camera with the [100] direction as precession axis. Weissenberg photographs were also taken about both the [100] and [110] directions as rotation axes. Mo  $K\alpha$  radiation was used throughout.

The diffraction symmetry was  $m3m$ . There were no systematic absences; thus the probable space groups are  $O_h^1-Pm3m$ ,  $T_d^1-P\bar{4}3m$  and  $O_1-P432$ .

The  $hkl$  and  $hkl$  intensity data used in the structure refinement were obtained from the precession photographs; the  $hkk$  data used in the refinement were obtained from the Weissenberg photographs, only those reflections for which  $\sin \theta_{hkk} < 0.790$  being included. The intensity estimates were made visually by comparison with a calibrated intensity strip. From the precession-camera data, relative  $|F_{hkl}|^2$  were obtained by applying the Lorentz-polarization corrections as obtained from Waser (1951, 1952) and Grenville-Wells & Abrahams (1952) charts supplied by Nies.† Lorentz-polarization corrections were also applied to the Weissenberg data in the usual way. No corrections were made for absorption or extinction.

### Refinement of the structure

It has been shown in the paper on  $\text{Pd}_{17}\text{Se}_{15}$ , that it is not possible to make a choice of the *most* probable of the three space groups to which the crystal might belong. The lowest standard errors were obtained in the centrosymmetric space group, but this in itself does not necessarily mean that it is the most probable one. In fact, as a result of the work on the  $\text{Pd}_{17}\text{Se}_{15}$ , the author is led to the speculation that if a structure truly belongs to a non-centrosymmetric space group but is only a very small perturbation on a centrosymmetric one, it is highly probable that it will converge more rapidly in the latter and have lower standard errors on the average. That is, the standard errors for like parameters will be about the same in the two groups, but those which arise from the additional degrees of freedom in the non-centrosymmetric group will have large standard errors associated with them (Geller, 1962). For  $\text{Rh}_{17}\text{S}_{15}$ , refinement was based only on the centrosymmetric space group  $Pm3m$ .

\* In some of the photographs, some extra spots appeared which were obviously from an occluded crystallite.

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The total number of independent reflections within the precession camera range on the zeroth and first layer photographs about [100] and on the zeroth level Weissenberg photograph about [110] with  $\sin \theta_{hkk} < 0.790$  was 402. Of these, 128 were too weak to be observed.

Refinement was carried out using the IBM 704 Busing-Levy (1959) least-squares program adapted to the IBM 7090 with a compatibility program. At first all reflection amplitudes of observed reflections were given unit weight while those of reflections too weak to be observed were weighted zero. The starting positional parameters were the analogous final parameters of  $\text{Pd}_{17}\text{Se}_{15}$  except for the Rh atoms in  $24m$ . These were taken as  $x=0.360$  and  $z=0.140$  as against 0.352 and 0.150 respectively for  $\text{Pd}_{17}\text{Se}_{15}$ . The change was based on examination of a number of the amplitudes. Starting values of isotropic temperature factors were  $0.10 \text{ \AA}^2$  for all Rh atoms and  $0.25 \text{ \AA}^2$  for all S atoms. In the calculation of structure amplitudes, the atomic scattering factors used for Rh were those of Thomas & Umeda (1957) and for sulfur, those of Dawson (1960). The real part of the dispersion correction (Dauben & Templeton, 1955) was applied to the Rh scattering factors.

Although after the first least-squares cycle there was little change in positional parameters, there was some difficulty with the temperature factors as was also found for  $\text{Pd}_{17}\text{Se}_{15}$ . In early iterations several of the amplitudes from reflections which suffered heavily from extinction were given zero weight and in subsequent iterations those for which  $|AF| \geq 30$  were given zero weight; the changes in positional parameters were not significant. The temperature factors were more reasonable but several were still negative\* in the last cycle. Examination of the correlation matrix (Geller, 1961) indicates that interactions between the temperature factors and positional parameters are all small and therefore that we may have some confidence in the interatomic distances calculated from the latter. The final parameters and standard deviations based on the use of 219 amplitudes with unit weight are given in Table 1. Calculated and observed amplitudes with the positional parameters given in Table 1 but with values of  $B$ ,  $0.05 \text{ \AA}^2$  for all Rh and  $0.10 \text{ \AA}^2$  for all S atoms are listed in Table 2. The discrepancy factor including considera-

\* However, the standard errors (Table 1) indicate that none of these are significantly different from zero.

Table 1. *Final parameters and standard errors*

Atom	Position	$x$	$y$	$z$	$B$	$\sigma(x)$	$\sigma(z)$	$\sigma(B)$
Rh	$24m$	0.3564	0.3564	0.1435	0.01	0.0001	0.0002	0.02
S	$12j$	0.1696	0.1696	$\frac{1}{2}$	-0.01	0.0006		0.09
S	$12i$	0.2310	0.2310	0	0.12	0.0008		0.11
S	$6f$	$\frac{1}{2}$	$\frac{1}{2}$	0.2643	0.14		0.0016	0.15
Rh	$6e$	0	0	0.2388	-0.05		0.0005	0.07
Rh	$3d$	0	0	$\frac{1}{2}$	0.00			0.07
Rh	$1b$	$\frac{1}{2}$	$\frac{1}{2}$	$\frac{1}{2}$	-0.24			0.15

Table 2. Observed and calculated amplitudes

hk0		hk1		hkk		hkk		hkk		hkk		hkk			
h	k	F <sub>o</sub>	F <sub>c</sub>	h	k	F <sub>o</sub>	F <sub>c</sub>	h	k	F <sub>o</sub>	F <sub>c</sub>	h	k	F <sub>o</sub>	F <sub>c</sub>
1	1	33	-27	2	2	137	-110	2	0	36	-9	8	3	104	-82
2	1	59	-37	3	2	193	-170	3	0	214	-226	9	3	322	-292
3	1	220	239	4	2	103	-97	4	0	195	-203	10	3	89	80
4	1	283	326	5	2	48	-42	5	0	225	-241	11	3	378	-412
5	1	170	189	6	2	< 67	-10	6	0	445	747	12	3	94	96
6	1	130	-115	7	2	< 73	-30	7	0	184	-181	13	3	144	144
7	1	152	-139	8	2	< 77	-30	8	0	474	709	14	3	170	-153
8	1	< 62	42	9	2	81	-71	9	0	127	127	15	3	41	47
9	1	178	166	10	2	98	-86	10	0	262	-307	16	3	42	42
10	1	141	116	11	2	90	-66	11	0	151	166	17	3	308	-294
11	1	81	62	12	2	< 91	4	12	0	196	233	18	3	90	88
12	1	85	70	13	2	< 90	-1	13	0	< 54	28	19	3	135	-105
13	1	71	44	14	2	240	264	14	0	616	553	20	3	79	-62
14	1	58	-52	15	2	103	-97	15	0	136	-120	21	3	152	146
2	2	< 35	9	16	2	< 64	-9	16	0	83	63	2	4	62	48
3	2	135	-111	17	2	< 68	-11	17	0	222	216	3	4	247	266
4	2	254	313	18	2	378	-447	18	0	77	-86	4	4	41	-58
5	2	< 51	-4	19	2	101	78	19	0	< 64	15	5	4	155	150
6	2	101	-91	20	2	< 83	-6	20	0	275	293	6	4	372	435
7	2	< 59	25	21	2	149	-153	21	0	63	-47	7	4	123	-131
8	2	< 61	18	22	2	160	161	22	0	327	272	8	4	584	573
9	2	< 64	-20	23	2	< 91	9	2	1	40	-32	9	4	218	187
10	2	110	111	24	2	227	-206	3	1	324	-423	10	4	230	-193
11	2	65	-54	4	4	20	-21	4	1	218	230	11	4	124	98
12	2	110	104	5	4	126	126	5	1	207	-191	12	4	184	139
13	2	51	34	6	4	57	81	6	1	157	-147	13	4	< 67	33
3	3	82	-68	7	4	144	144	7	1	101	105	14	4	455	5
4	3	139	-124	8	4	121	122	8	1	56	48	15	4	< 71	-35
5	3	51	40	9	4	84	78	9	1	109	-102	16	4	85	71
6	3	< 85	-2	10	4	89	-99	10	1	58	41	17	4	202	157
7	3	185	190	11	4	72	-61	11	1	270	-288	18	4	110	-88
8	3	186	172	12	4	73	52	12	1	< 58	32	19	4	< 78	15
9	3	162	131	13	4	189	168	13	1	< 63	53	20	4	279	256
10	3	163	-167	14	4	74	-81	14	1	< 65	-12	21	4	< 76	6
11	3	144	-157	15	4	72	-52	15	1	< 70	14	2	5	< 42	-11
12	3	68	-59	16	4	199	-204	16	1	< 72	41	3	5	94	-88
13	3	204	174	17	4	134	128	17	1	155	-168	4	5	183	-73
4	4	573	1002	18	4	< 87	-5	18	1	< 77	24	5	5	77	-73
5	4	162	144	19	4	< 90	-23	19	1	110	-88	6	5	69	-54
6	4	207	-203	20	4	71	-71	20	1	< 78	-1	7	5	110	-17
7	4	245	239	21	4	< 91	34	21	1	63	86	8	5	145	134
8	4	< 62	0	22	4	93	-101	22	2	23	24	9	5	< 56	-2
9	4	96	101	23	4	142	-163	23	2	30	18	10	5	< 59	20
10	4	489	549	24	4	< 85	-32	24	2	< 34	1	11	5	119	-129
11	4	87	-89	25	4	< 89	-11	25	2	109	-107	12	5	53	-5
12	4	298	294	26	4	< 92	31	26	2	< 40	-20	13	5	< 68	-9
13	4	266	220	27	4	< 93	-30	27	2	43	-45	14	5	< 70	32
5	5	137	133	28	4	< 93	3	28	2	54	66	15	5	< 74	-54
6	5	< 59	25	29	4	191	212	29	2	41	-44	16	5	< 76	48
7	5	61	55	30	4	< 93	-32	30	2	54	-31	17	5	98	-38
8	5	191	156	31	4	155	-151	31	2	< 59	-18	18	5	< 79	21
9	5	141	150	32	4	< 73	44	32	2	< 61	-2	19	5	< 79	-67
10	5	< 65	-39	33	4	366	-334	33	2	63	-70	20	5	< 77	29
11	5	< 61	-39	34	4	< 97	5	34	2	< 67	16	21	5	77	-21
12	5	63	67	35	4	152	< 70	35	2	70	4	2	6	62	-51
6	6	293	351	36	4	162	29	36	2	72	29	4	6	211	-182
7	6	148	-164	37	4	< 93	88	37	2	< 76	-46	5	6	51	-37
8	6	283	320	38	4	< 93	83	38	2	< 77	-49	6	6	204	-196
9	6	65	-46	39	4	72	62	39	2	< 79	-25	7	6	90	-94
10	6	213	-246	40	4	< 92	-42	40	2	< 79	45	8	6	128	131
11	6	60	53	41	4	< 77	44	41	2	< 77	-15	9	6	109	-106
12	6	101	106	42	4	101	106	42	3	59	-42	10	6	227	-178
7	7	215	-219	43	4	862	-813	43	3	662	-613	11	6	< 66	40
8	7	65	-53	44	4	200	202	44	3	200	202	12	6	67	53
9	7	83	93	45	4	251	-232	45	3	251	-232	13	6	108	-84
10	7	106	126	46	4	77	-90	46	3	77	-90	14	6	146	-116
11	7	< 51	36	47	4	330	379	47	3	330	379	15	6	106	-92
12	7	365	407	48	4	101	106	48	3	101	106	16	6	146	-116
9	8	114	141	49	4	101	106	49	3	101	106	17	6	146	-116
10	8	106	-131	50	4	101	106	50	3	101	106	18	6	146	-116
11	8	97	101	51	4	101	106	51	3	101	106	19	6	146	-116
9	9	230	223	52	4	101	106	52	3	101	106	20	6	146	-116
10	9	< 54	-21	53	4	101	106	53	3	101	106	21	6	146	-116

tion of multiplicity but excluding the 600, 800 and 440 amplitudes is 0.12. When the three amplitudes are included, the discrepancy factor is 0.13.

### Description of the structure

The interatomic distances in  $Rh_{17}S_{15}$  are listed in Table 3. For ease of comparison the analogous  $Pd_{17}Se_{15}$  distances are also listed. The general features of the  $Rh_{17}S_{15}$  structure are shown in Fig. 2 of the paper on  $Pd_{17}Se_{15}$  (Geller, 1962). However, in detail, there are important differences in the  $Rh_{17}S_{15}$  and  $Pd_{17}Se_{15}$  structures.

The  $Rh(b)$ - $S(f)$  are shorter than the  $Rh(d)$ - $S(j)$  distances (Table 3), even though the former are octahedral whereas the latter appear to be square-coordination distances. This is reverse of the situation in  $Pd_{17}Se_{15}$ . On the other hand,  $Rh(d)$  also has two  $Rh(e)$  neighbors at 2.59 Å which together with the 4  $S(j)$  atoms completes an octahedron about  $Rh(d)$ . In  $Pd_{17}Se_{15}$ , the analogous  $Pd(d)$ - $Pd(e)$  distance is

2.78 Å, which is 0.03 Å longer than the elementary Pd-Pd distance, 2.75 Å. In elementary Rh which has the same cubic close-packed structure as elementary Pd, the nearest neighbor Rh-Rh distance is 2.69 Å. Thus the  $Rh(d)$ - $Rh(e)$  distance in  $Rh_{17}S_{15}$  is 0.10 Å shorter than the elementary Rh-Rh distance, indicating considerably stronger bonding than between the analogous  $Pd(d)$ - $Pd(e)$  neighbors. This would account for the longer  $Rh(d)$ - $S(j)$  distances.

Further corroboration of the preceding argument is obtained from examination of the Rh-S distances involving  $Rh(e)$ , which is surrounded by a square of  $S(i)$  atoms (though  $Rh(e)$  is 0.08 Å out of the plane of the  $S(i)$  atoms) and one  $Rh(d)$  atom. The  $Rh(e)$ - $S(i)$  distance, 2.29 Å, is shorter than that of  $Rh(b)$ - $S(f)$ , but not in the same proportion as the analogous distances in  $Pd_{17}Se_{15}$ . Again this is justified by the shorter  $Rh(e)$ - $Rh(d)$  distance.

The average  $Rh(m)$ - $S$  distance is 2.33 Å, not significantly different from the  $Rh(b)$ - $S(f)$  distance. This appears to be somewhat anomalous, but again

Table 3. *Interatomic distances and standard error in Rh<sub>17</sub>S<sub>15</sub>*

Atom	Neighbors	Distance	$\sigma$	Analogous Pd <sub>17</sub> Se <sub>15</sub> distance
Rh( <i>b</i> )	6 S( <i>f</i> )	2.336 Å	0.016 Å	2.576 Å
Rh( <i>d</i> )	4 S( <i>j</i> )	2.378	0.009	2.527
	2 Rh( <i>e</i> )	2.588	0.007	2.781
Rh( <i>e</i> )	4 S( <i>i</i> )	2.291	0.008	2.438
	1 Rh( <i>d</i> )	2.588	0.007	2.781
	4 Rh( <i>e</i> )	3.348	0.007	3.566
Rh( <i>m</i> )	1 S( <i>i</i> )	2.261	0.010	2.430
	1 S( <i>f</i> )	2.342	0.010	2.492
	2 S( <i>j</i> )	2.349	0.006	2.509
	1 Rh( <i>m</i> )	2.844	0.004	3.184
	2 Rh( <i>m</i> )	2.846	0.003	3.137
	2 Rh( <i>m</i> )	2.984	0.003	3.029
S( <i>f</i> )	4 S( <i>f</i> )	3.304	0.022	3.643
	4 S( <i>j</i> )	3.406	0.012	3.641
	4 Rh( <i>m</i> )	2.342	0.010	2.492
	1 Rh( <i>b</i> )	2.336	0.016	2.576
S( <i>i</i> )	4 S( <i>i</i> )	3.238	0.009	3.445
	4 S( <i>j</i> )	3.210	0.007	3.440
	2 Rh( <i>e</i> )	2.291	0.008	2.438
	2 Rh( <i>m</i> )	2.261	0.010	2.430
S( <i>j</i> )	2 S( <i>f</i> )	3.406	0.012	3.641
	4 S( <i>i</i> )	3.210	0.007	3.440
	2 S( <i>j</i> )	3.362	0.013	3.572
	1 Rh( <i>d</i> )	2.378	0.009	2.527
	4 Rh( <i>m</i> )	2.349	0.006	2.509

examination of the metal-metal distances indicates a possible reason. The three nearest Rh(*m*)-Rh(*m*) distances, although 0.16 Å longer than the elementary distances, are relatively considerably shorter than the analogous Pd-Pd distances; the nearest neighbor Pd(*m*)-Pd(*m*) distances average 3.15 Å which is 0.40 Å longer than the elementary distance. Thus it appears that in Rh<sub>17</sub>S<sub>15</sub> the Rh-Rh bonds are in general stronger than the Pd-Pd bonds in Pd<sub>17</sub>Se<sub>15</sub>.

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## Neutron Diffraction Study of Ammonium Perchlorate

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A neutron diffraction study of single crystals of NH<sub>4</sub>ClO<sub>4</sub> at room temperature has indicated that the orientation of the ammonium ion in the crystal is random, and suggests that it is undergoing free or nearly free rotation.

### Introduction

A room temperature single crystal study of ammonium perchlorate was undertaken in order to compare the structure of this compound with the 'isomorphous' compound hydronium perchlorate (also under investigation by the authors). A structure comprising an ordered, hydrogen-bonded arrangement of the ammonium ions had been suggested from X-ray

difference Fourier projections by Venkatesan (1957). This arrangement, however, is not consistent with the present neutron data which suggests that the ammonium ion is undergoing essentially free rotation. The recent infra-red (Waddington, 1958), nuclear magnetic resonance (*NMR*) (Ibers, 1960, and Richards & Schaefer, 1961), heat capacity (Justice & Westrum, 1961) and cold neutron (Rush, Taylor & Havens, 1961) studies strongly support these results. The OH<sub>3</sub>ClO<sub>4</sub> neutron diffraction study in which the OH<sub>3</sub><sup>+</sup> ion is disordered, but not rotating, will be reported at a later date.

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